

## TEPHRA ATTRIBUTION

**Sample SV 45.40** - Glasses of juvenile fragments extracted from SV 45.40 define two clusters straddling the boundary between trachytes and phonolites. The first groups glasses with a SiO<sub>2</sub> content ca. 61 wt% and a K<sub>2</sub>O/Na<sub>2</sub>O ratio slightly higher than 1, a composition which is attributable to the Campanian Ignimbrite pyroclastic fall deposits. The second cluster has SiO<sub>2</sub> wt% varying between 61.7 and 57.6 and a K<sub>2</sub>O/Na<sub>2</sub>O ratio exceeding 2, which is a typical feature of glasses from the Pomici di Base (Somma-Vesuvius) eruption, which constraints this layer as younger than ca. 22 ka (Santacroce et al., 2008).

**Samples SV 44.70 and SV 41.10** - The glasses extracted from the juvenile fragments that make up these layers belong to three chemical clusters, two of which are completely identical to those of the previous sample, and the third is phonolitic with SiO<sub>2</sub> around 59 wt% and a K<sub>2</sub>O/Na<sub>2</sub>O ratio of around 0.9, which are typical characteristics of glasses from the Mercato eruption (Somma-Vesuvius), the presence of which limits the age of the layer to less than 9 ka (Santacroce et al., 2008).

**Sample SV 35.40** - The glasses of the juvenile fragments are grouped into three distinct clusters, one belonging to the trachyte fields, another along the boundary between trachyte and phonolite, and a third belonging to the phonolite field. These have been attributed to the Pomici di Base (Somma-Vesuvius; ca. 22 ka), Monte Spina (Campi Flegrei; ca. 4.5 ka, Smith et al., 2011) and Avellino (Somma-Vesuvius; 3870 cal yrs BP, Passariello et al., 2009) eruptions, respectively, which is the age beyond which this layer cannot be older.

**Sample SV 33.50** - The glasses of all juvenile fragments extracted from this layer are phonolitic in composition (Fig. 8a<sub>1</sub>,a<sub>2</sub>) and well correspond to the composition of the protohistoric AP2 (Somma-Vesuvius) eruption (Rolandi et al., 1998; Santacroce et al., 2008), dated to 3380±23 y (Passariello et al., 2009), which corresponds to 3615 cal yrs BP.

**Sample SV 31.80** - Glasses of juvenile fragments extracted from SV 31.80 define two clusters straddling the boundary between trachyte and phonolite, allowing us to identify the presence of both CI and Pomici di Base fragments (Fig. 8b<sub>1</sub>, b<sub>2</sub>). A third smaller cluster straddles the boundary between phonolites and tephriphonolites and can be related to the glasses of the protohistoric AP4 (Somma-Vesuvius) eruption, for which an exact age is not available in the literature. According to Rolandi et al. (1998), it could be younger than 2800 yrs cal BP,

but many marine age models suggest that the AP3 eruption, the protohistoric eruption which precedes AP4 found in the FSV sequence, should have occurred around 3.4 ka (see Totaro et al., 2022 and references therein).

**Sample SV 27.70** - Glasses of this layer have the same features of those of sample SV 45.40, and are ascribed to Campanian Ignimbrite (Campi Flegrei) and Pomici di Base (Somma-Vesuvius) eruptions, giving no useful age constraint.

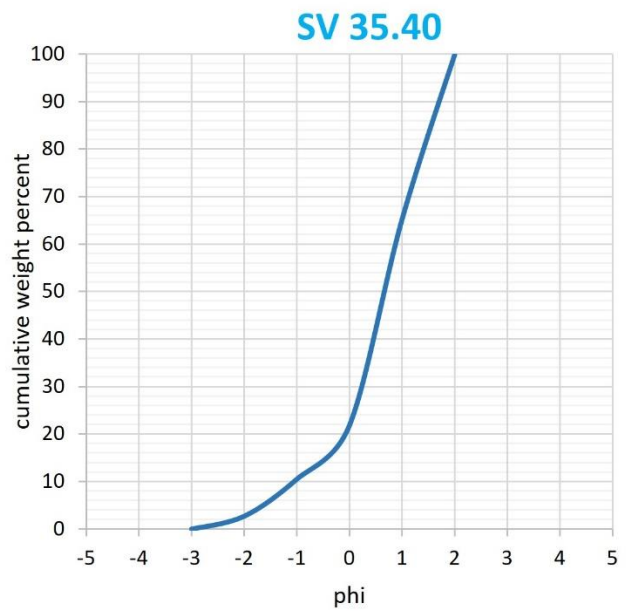
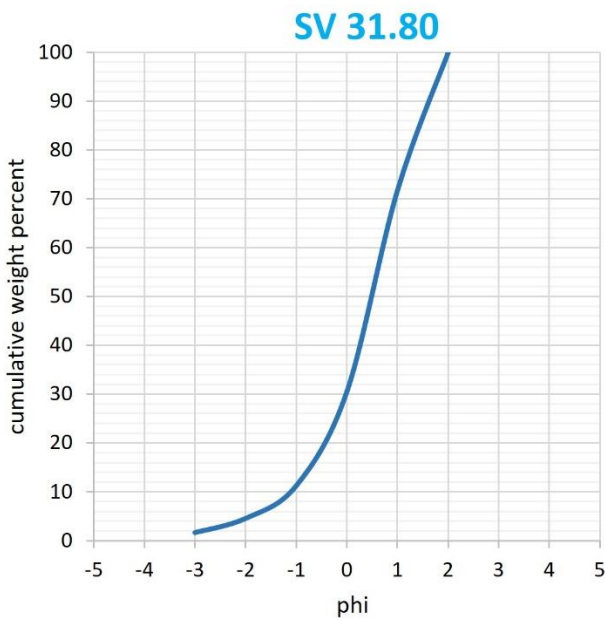
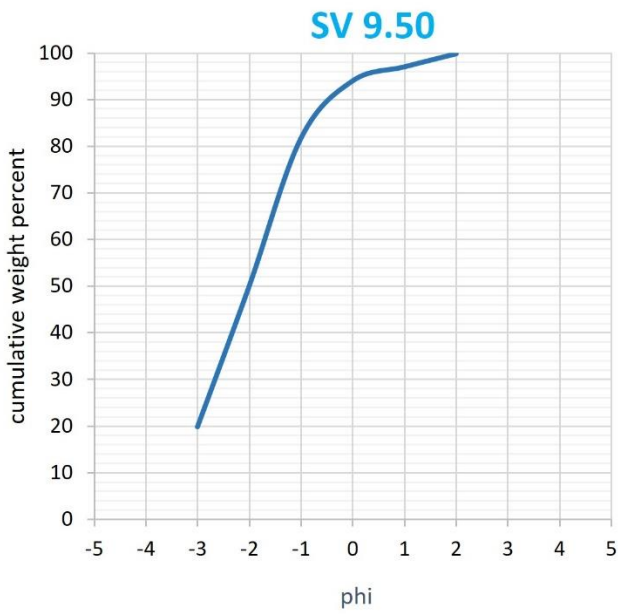
**Sample SV 26.10** - The glasses of juvenile fragments making up this layer are very variable in composition, since they define two clusters in the trachyte field, distinguishable on the basis of the  $K_2O/Na_2O$  ratio and on the  $SiO_2$  wt% content and corresponding, respectively, to Campanian Ignimbrite and Monte Spina eruptions (both from Campi Flegrei). Other two clusters fall in the phonolite field, distinguishable on the basis of the  $Na_2O$  content, which is higher (ca. 9 wt%) for the almost aphyric pumices of the Avellino eruption (Somma-Vesuvius), and lower (ca. 5.5 wt%) for the leucite bearing CE 79 (Somma-Vesuvius) pumice fragments, which, according to our first appearance approach, implies that this level could not be older than CE 79.

**Sample SV 25.50** - Two phonolitic clusters group the juvenile fragments of this layer, fully comparable with those of sample SV 26.10.

**Sample SV 22.50** - All the juvenile fragments that make up this primary tephra layer are foiditic to phonotephritic in composition and well comparable with the composition of the juvenile fragments of the CE 472 (Pollena) eruption from Somma-Vesuvius (Santacroce et al., 2008; Melluso et al., 2022).

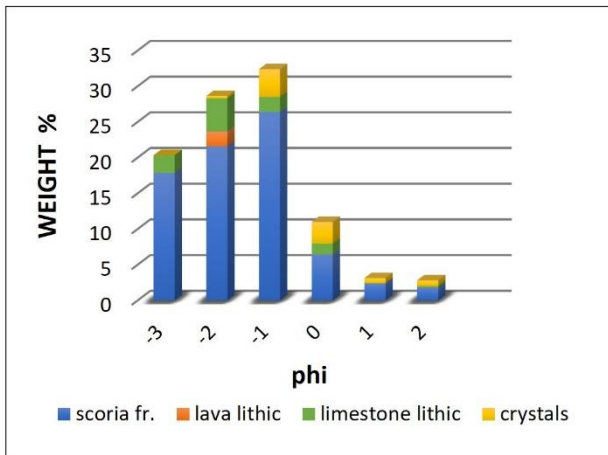
**Sample SV 9.50** - The glasses from juvenile fragments of this primary tephra layer are low-silica phonolites (Fig. 8c<sub>1</sub>,c<sub>2</sub>) and well resemble the composition of the glasses extracted from the Somma- Vesuvius 1631 eruption (Santacroce et al., 2008).

**Fig. 1SM** - Grain-size cumulative plots of the fraction coarser than 250 micron for the primary samples SV 9.30 and SV 22.10, and for the reworked samples SV 31.80 and SV 35.40

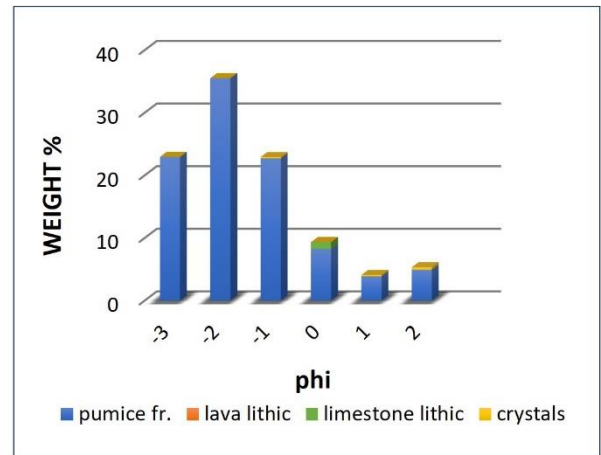


**Fig. 2SM** - Lithological component analysis of the fraction coarser than 250 micron for the primary samples SV 9.30 and SV 22.10, and for the reworked samples SV 31.80 and SV 35.40

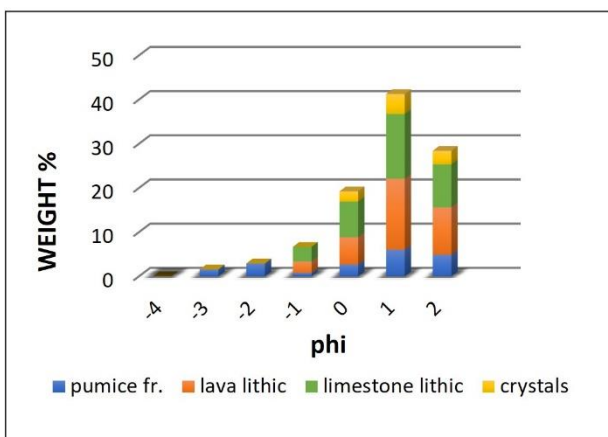
**SV 9.30**



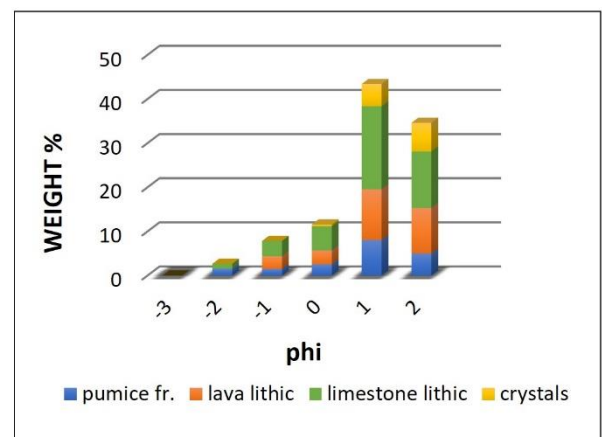
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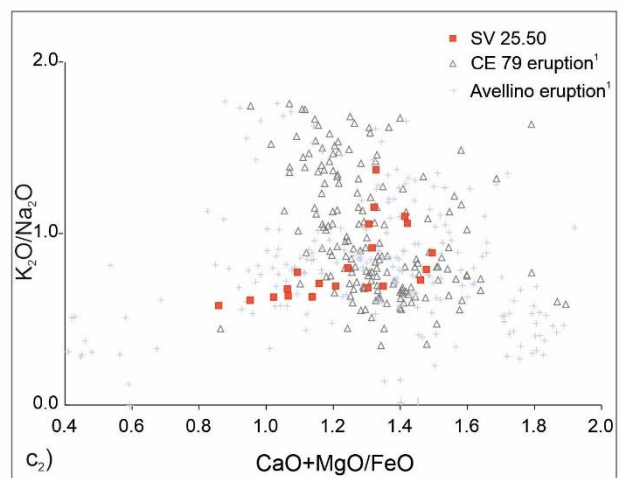
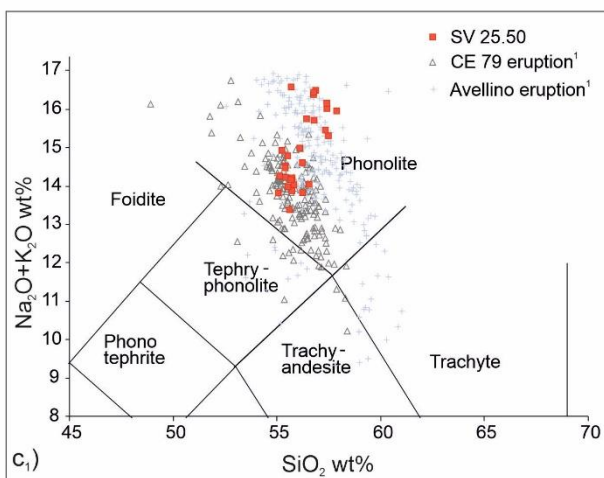
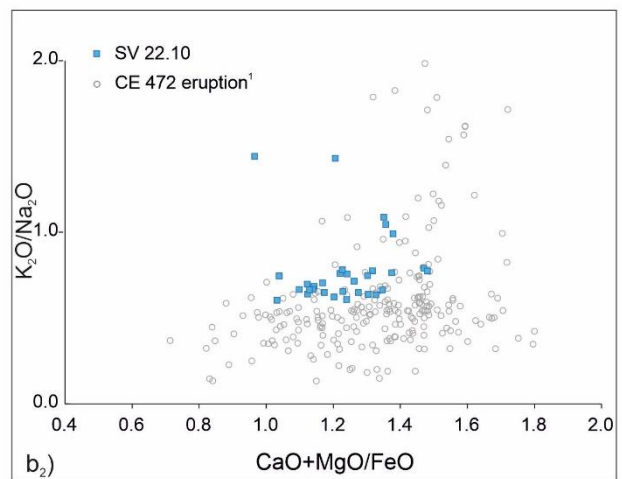
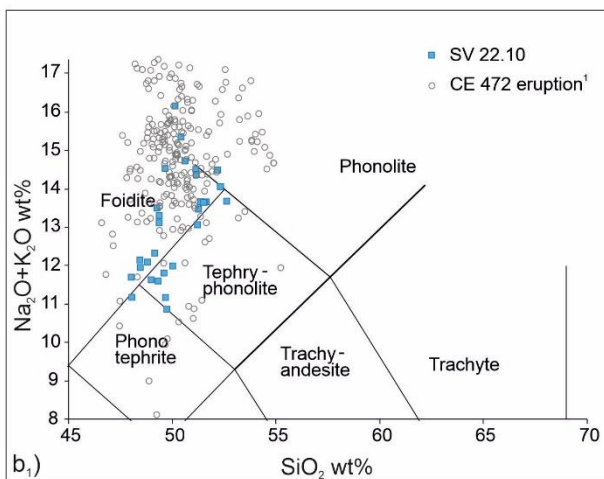
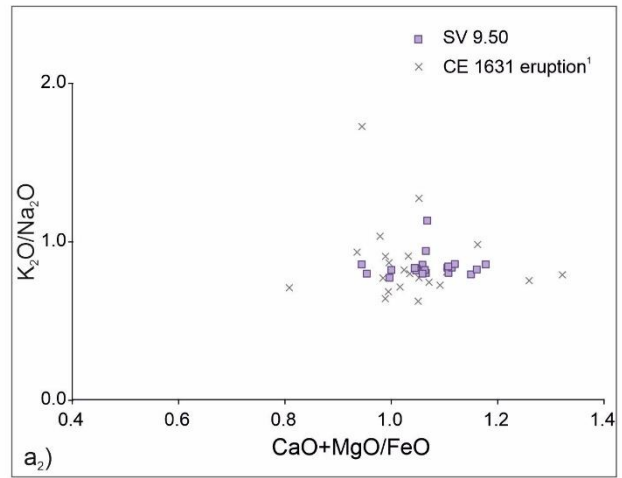
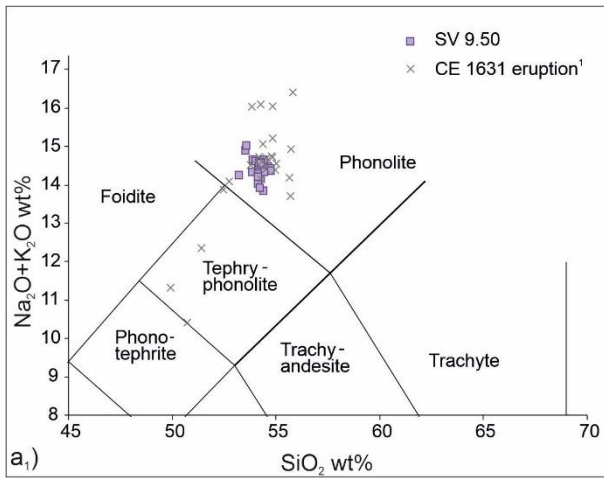
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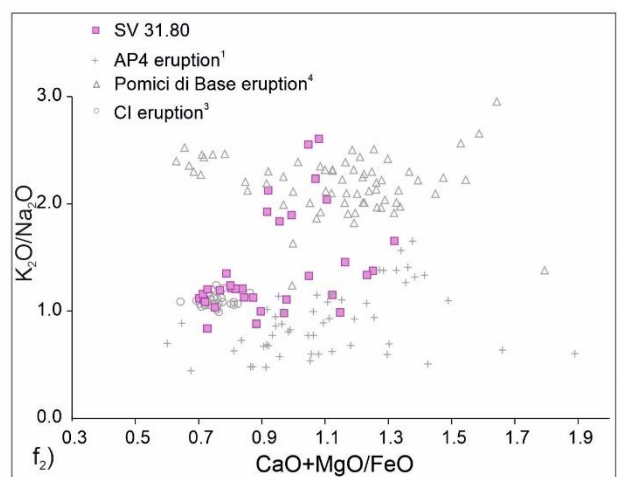
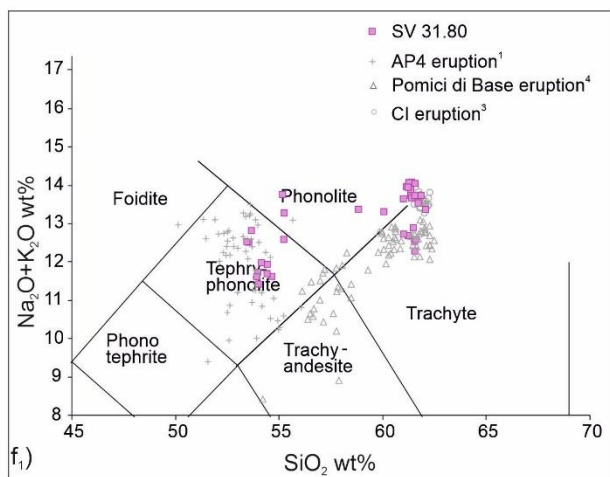
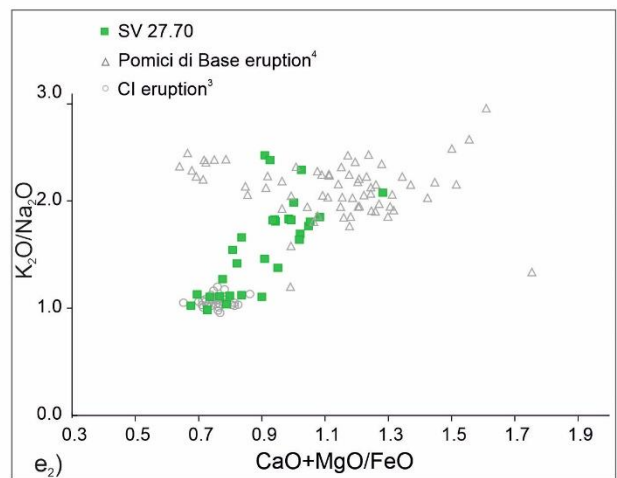
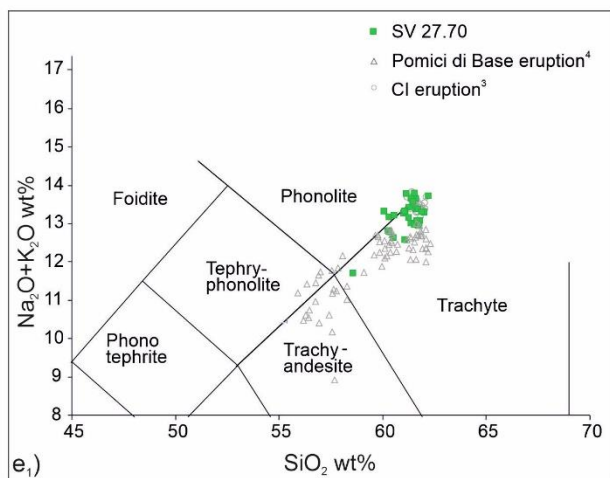
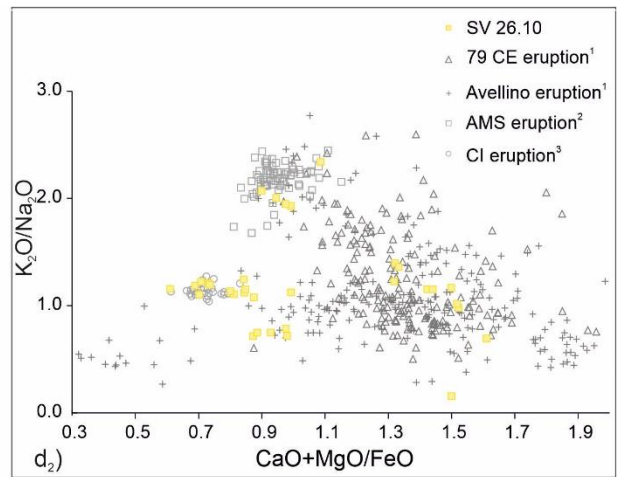
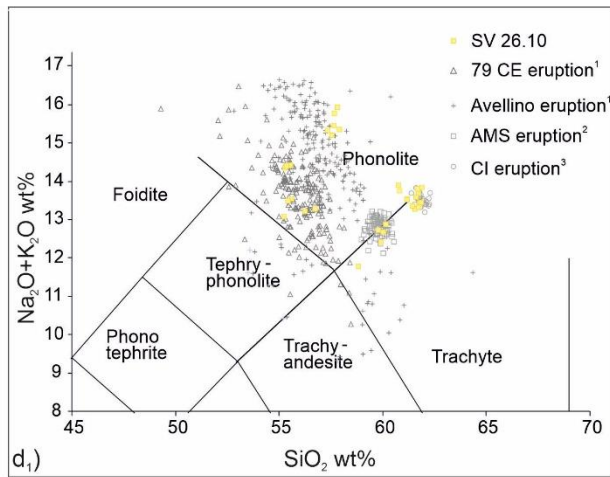


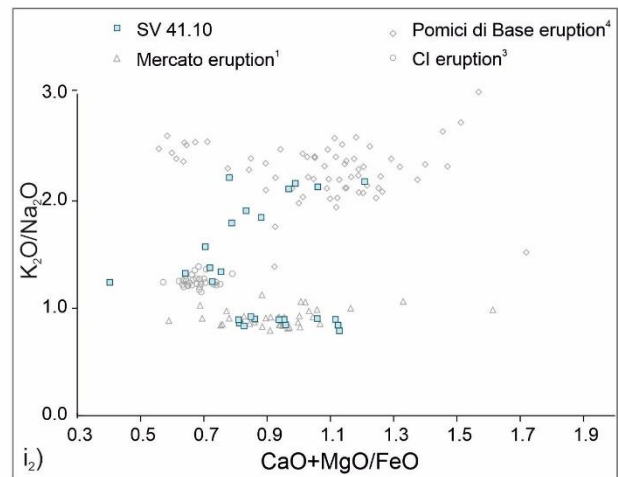
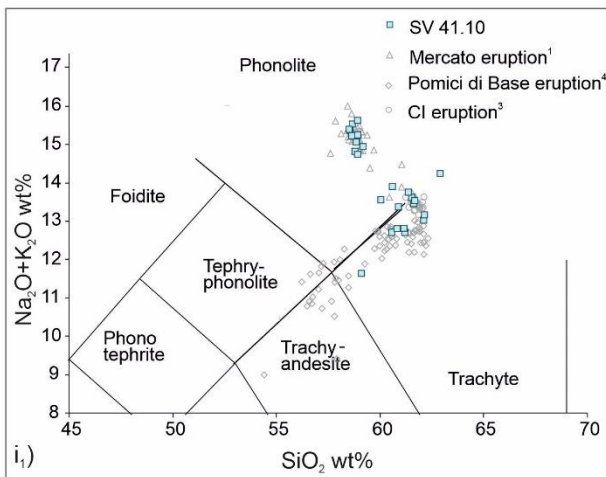
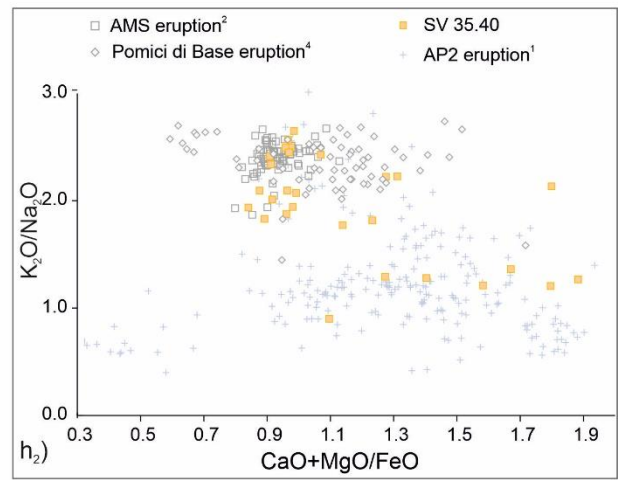
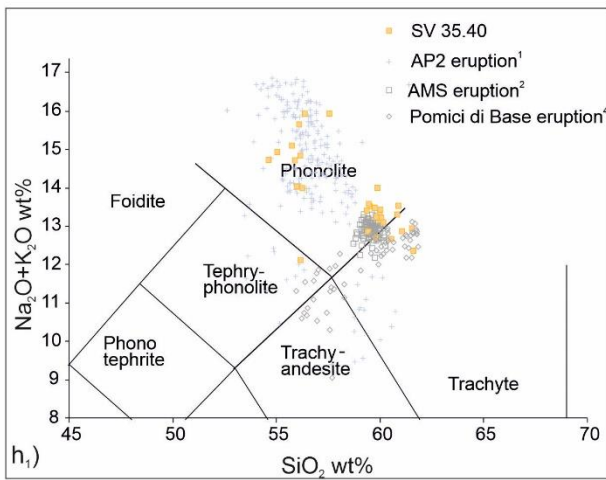
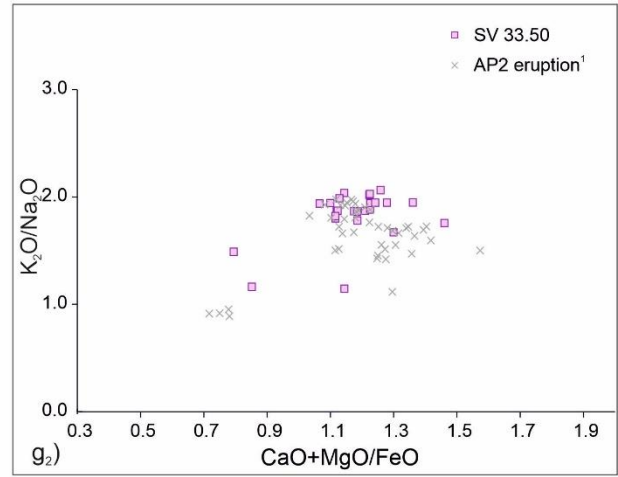
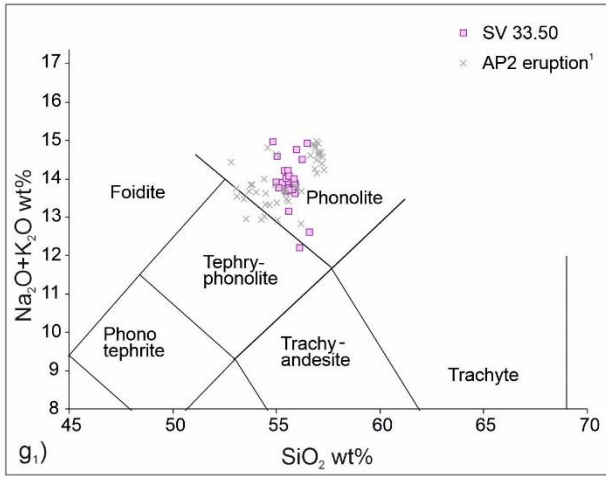
**SV 35.40**



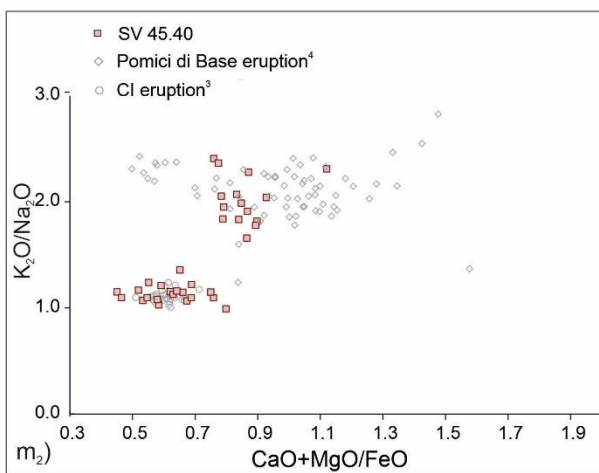
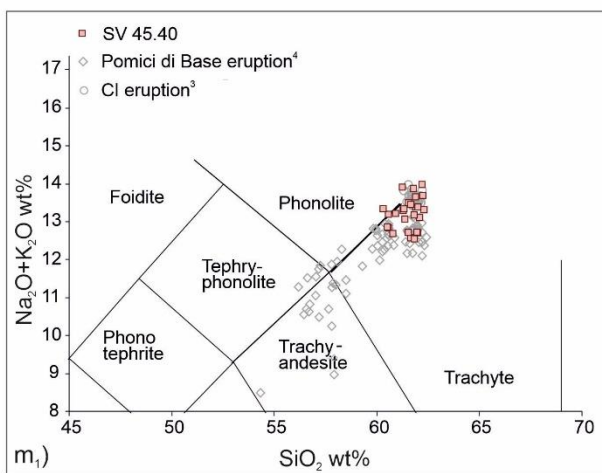
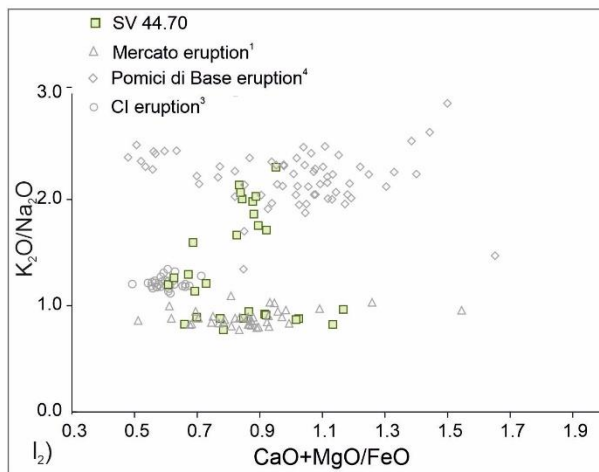
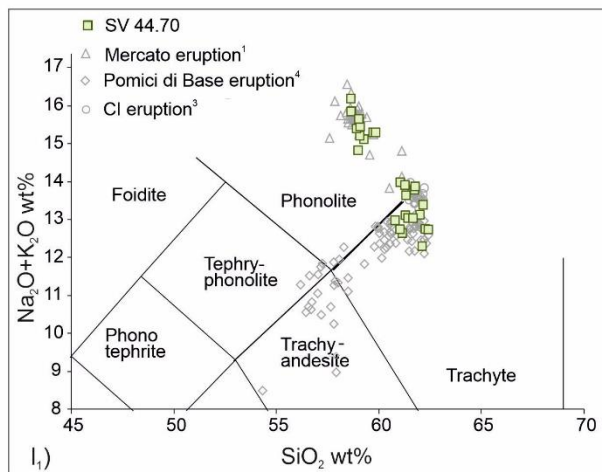
**Fig. 3SM** – Total Alkali – silica plots ( $a_1, b_1, c_1$ , etc.) and  $\text{CaO}+\text{MgO}/\text{FeO}$  vs  $\text{K}_2\text{O}/\text{Na}_2\text{O}$  plots ( $a_2, b_2, c_2$ , etc.) for glasses extracted from juvenile fragments of the investigated tephra layers compared to the products of the correlative eruptions. <sup>(1)</sup> data from Santacroce et al. (2008); <sup>(2)</sup> data from Smith et al. (20112); <sup>(3)</sup> data from Tomlinson et al. (2012); <sup>(4)</sup> data from Pappalardo et al. (2018).













## References

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